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University of Rhode Island

Telecon ESIP Information Quality Cluster

9 October 2018

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Outline



SST Error Budget

- Constraints on the Error Budget
- Overview
- Products
- Data levels and processing steps
- The error budget
- Two Take-Aways

Determining SST & VIIRS Instrument Noise

- Introduction
- Two Approaches to Determining the Instrument Noise

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- Data Preparation
- Results

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 - O The physical basis of SST measurements;
 - Radiative transfer modeling and SST retrieval algorithm development;
 - O Cal/val pre-launch and on-orbit:
 - Data merging and gridding:
 - O The climate record: representation data access and stability and:
 - O Applications of SST6
- Groups representing each area assembled their findings in a report.
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BUT - There is little that ties this error budget to SST

Steering Committee

The Steering Committee for the workshop and subsequent White Paper:

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- Sandra Castro (U Colorado)
- Peter Cornillon (U Rhode Island) that'd be me.
- Chelle Gentemann (Remote Sensing Systems, Inc)
- Peter Hacker (NASA)
- Andy Jessup (U Washington)
- Alexey Kaplan (Columbia U)
- Eric Lindstrom (NASA)
- Eileen Maturi (NOAA)
- Peter Minnett (U Miami)
- Dick Reynolds (Coop. Inst. for Climate and Sat.)

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Constraints on the SST Error

- The Applications Group identified acceptable bounds on SST products
 - Spatial resolution,
 - Temporal resolution,
 - Geolocation accuracy,
 - Absolute SST accuracy, and
 - Relative SST accuracy.
- These bounds were categorized by application.

Applications	Source	Spatial	Temporal	Geolocation	Absolute	Relative
		resolution	resolution	accuracy	accuracy	accuracy
		(km)	(hrs)	(km)	(^o K)	
CDR	Ohring et al., 2005				0.1	0.04 [°] K/decade
CDR	Workshop					0.05 [°] K/decade
NWP	Eyre et al., 2009	5	3		0.3	
Global Operations	NPOESS IORD-II	0.25	3	0.1	0.1	0.05 ⁰ K
Coastal/Lake Operations	NPOESS IORD-II	0.1	6	0.1	0.1	
Fronts	Workshop	0.1	0.25	0.1	1	0.1 [°] K
Climate Models	Workshop	25	24	5	0.2	0.05 [°] K/decade
Lakes	Workshop	1	3	1	0.3	0.2 ⁰ K
Air-sea Fluxes	Workshop	10	24	2	0.1	
Mesoscale	Workshop	1	168		0.1	
Submesoscale	Workshop	0.1	1		0.1	
Strictest		0.1	0.25	0.1	0.1	0.05 [°] K 0.04 [°] K/decade



A significant fraction of workshop participants were interested in feature studies.

Such studies tend to be underrepresented in specification of product uncertainty.

Different uses place different demands on the characteristics of the error that are of interest



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Approach to Developing the Error Budget

The error budget is discussed in the white paper from 2 perspectives

- Two groups of products
- Five NASA Product levels
- Although both approaches were considered

The focus was on product categories

Constraints Overview Products Levels Errors Take-aways

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Skin/subskin SST Retrievals in Satellite Coordinates

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 - Subskin products are
 - Derived from microwave sensors.
 - Obtained by combining radiances in different spectral bands with slightly differences sense looks into
 - In both cases
 - The primary conversion is radiance to SST.
 - Upper x1 mm
- Derived SST products
 - Products interned from skin/subskin SST retrievals.
 - These products generally require:

 Requires assumptions about spatial and temporal variability of temperatures in the upper costan.

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Derived SST products

- These products generally require:
 - Regridding and/or
 - Collating and/or
 - Adjustment to a depth below 1mm and/or
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 - Adjustment to a depth below 1mm and/or
 - Interpolation into gaps.
- Requires assumptions about spatial and temporal variability of temperature in the upper ocean.

Constraints Overview Products Levels Errors Take-aways

Data Products Schematically



- Satellite-derived data products are generally divided into 5 categories.
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Data Level Perspective

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Data levels and processing schematically


Errors associated with skin/subskin retrievals fall in two groups

- Instrument error
- Retrieval algorithm error
- Errors for derived products also fall into two groups
- Errors introduced at any step propagate to the next step.

So let's look at these errors in more detail



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- Instrument noise
- Calibration source
- Characterization of the instrument
- Stray radiation
- Location of the observation



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- Input uncertainties in ancillary data used by the geophysical model(s)
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- Oceanographic variability gives rise to errors resulting from:
 - Temporal changes in SST when combining skin/subskin values over time
 - Spatial differences when estimating temperatures at different depths

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Merging, Gridding and Analysis Errors



• Merging, gridding and analysis errors result from:

- The procedure used to merge values from different sensor/passes.
- Biases in the data from one source relative to another.
- Differences between the input and output grids.
- Method used to interpolate to locations for which there are no SST retrievals
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Take-Aways

- This approach is readily generalizable to most other satellite-derived parameters of interest to the Earth science community
- It is not just the absolute SST error that is important:
 - Location error;
 - Pixel characterization.
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Outline

Background

2 SST Error Budget

- Constraints on the Error Budget
- Overview
- Products
- Data levels and processing steps
- The error budget
- Two Take-Aways

Determining SST & VIIRS Instrument Noise

- Introduction
- Two Approaches to Determining the Instrument Noise

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- Data Preparation
- Results

Overview - Measures of SST Uncertainty in Satellite-Derived Fields

• The uncertainty of satellite SST data products is determined from in situ matchups.

- Standard measure is rms difference between buoy and satellite SSTs.
- Typical values for AVHRR, MODIS . . . range from 0.4 to 0.6 K.
- But these are based on match-ups widely separated in space and time
 - A significant contributor to these uncertainties are atmospheric fluctuations
 - Which vary over large scales...
- But I'm interested in SST fronts and gradients,
 - For which large scale variability is relatively unimportant.
 - I want to know the pixel-to-pixel noise and such measures are not available.

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- 4-km global Pathfinder SST fields (AVHRR).
- 4-km global MODIS SST fields.
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- The grid onto which they were projected was identical.
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Accuracy versus Precision and the Local Precision



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Back to the SST error budget.

Error Budget for Satellite-Derived SST Fields (NASA SST Science Team)



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Two Approaches to Determining 'Instrument Noise'

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The Spectral Approach

- Wavenumber spectrum in the Sargasso Sea at scales larger than 1 km is very nearly linear in log-log space.
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 - Along-scan and along-track characteristics differ.
- And these groups were further subdivided into day and night subgroups.
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Step 1. Generate the Spectra

- The sections were nearest neighbor resampled to equal spacing.
- The evenly spaced, complete temperature sections were demeaned.
- Then Fourier Transformed NO filtering
 - Filtering impacts high wavenumber portion of the spectrum.
- Spectra ensemble averaged

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Step 2. Fit Straight Line + Noise to Mean Spectra

Straight line spectra plus white noise were fit to mean spectra.

 $\text{PSD}_{\text{Fit}} = 10^{(\textit{log}_{10}(\text{Wavenumber})*\text{Slope}+\text{Intercept})} + \text{Noise}$

• Where we minimize:

 $\xi = \left(\log_{10}(extsf{PSD}_{ extsf{Fit}}) - \log_{10}(extsf{PSD}_{ extsf{Obs}})
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Sample Results for AVHRR Along-Scan



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Sample Results for VIIRS Along-Scan



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AVHRR and VIIRS Nighttime, Along-Scan Compared



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Results

	Method	Day (K)		Night (K)	
		Along-Scan	Along-Track	Along-Scan	Along-Track
AVHRR	Spectral	0.172	0.209	0.173	0.209
	Variogram	0.185	0.219	0.183	0.219
VIIRS	Spectral	0.046	0.076	0.021	0.032
	Variogram	0.081	0.097	0.042	0.056

Summary

- Estimates for cloud-free regions *instrument* noise only; no *classification* error.
- Variogram estimates slightly larger than spectral estimates for AVHRR; but track well.
- Variogram estimates about twice spectral estimates for VIIRS.
 - Likely due to $\sigma_{inst} \approx \sigma_{geo}$.
- Along-Track noise > Along-Scan noise.; \approx 1.5 \times for VIIRS
- Daytime noise > Nighttime noise; $\approx 2 \times$ for VIIRS
- AVHRR results for NOAA-15, other AVHRRs may be less noisy;.

NOAA-15 Noise versus Time



A Condsequence

How does noise impact satellite-derived SST gradients?

To determine this we

- Simulated 10,000 3 × 3 pixel squares for a fixed gradient in x, $\frac{\partial T}{\partial x}$, 0 in y.
- Added Gaussian white noise, σ , to each of the elements.
- Applied the 3×3 Sobel gradient operator in x and y.
- Determined the μ and σ of the resulting gradient components and the gradient magnitude.
- Performed the above for:

$$0.01 \text{ K km}^{-1} < \frac{\partial T}{\partial x} < 0.3 \text{ K km}^{-1}$$
$$0.001 \text{ K} < \sigma < 0.3 \text{ K}$$

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A Condsequence – Gradient Components



A Condsequence – Gradient Magnitude



Numerous authors have published gradient magnitude fields from AVHRR Including me – GULP!

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A Condsequence – Gradient Magnitude



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